

Body	Body Symbol	μ -GM (km ³ /s ²)	SOL radius (km)
Sun		1.32712440017987 × 10 ¹¹	—
Mercury		2.2032080486 × 10 ⁴	112 000
Venus		3.24858598826 × 10 ⁵	616 000
Earth		3.98600435608 × 10 ⁵	925 000
(Moon)		4.902799108 × 10 ³	66 100
Mars*		4.2828314258 × 10 ⁴	577 000
Jupiter*		1.2671276785779 6 × 10 ⁸	48 200 000
Saturn*		3.7940626061137 × 10 ⁷	54 800 000
Uranus*		5.794549007072 × 10 ⁶	51 800 000
Neptune*		6.836534063879 × 10 ⁶	86 600 000
(Pluto)*		9.81600888 × 10 ²	3 080 000

*Includes the gravitational parameters of the planet's satellites

Gravitational parameters

Object	Radius (km)	Mass (kg)	Sidereal Rotation Equator to Orbit Plane	Inclination of Orbit	Semimajor Axis of Orbit (km)	Orbit Eccentricity	Inclination of Orbit to the Ecliptic Plane	Orbit Sidereal Period
Sun	696000	1.989 × 10 ³⁰	25.38d	7.25°	—	—	—	—
Mercury	2440	330.2 × 10 ²¹	58.65d	0.01°	57.91 × 10 ⁶	0.2056	7.00°	87.97d
Venus	6052	4.869 × 10 ²⁴	243d*	177.4°	108.2 × 10 ⁶	0.0067	3.39°	224.7d
Earth	6378	5.974 × 10 ²⁴	23.9345h	23.45°	149.6 × 10 ⁶	0.0167	0.00°	365.256d
(Moon)	1737	73.48 × 10 ²¹	27.32d	6.68°	384.4 × 10 ³	0.0549	5.145°	27.322d
Mars	3396	641.9 × 10 ²¹	24.62h	25.19°	227.9 × 10 ⁶	0.0935	1.850°	1.881y
Jupiter	71490	1.899 × 10 ²⁷	9.925h	3.13°	778.6 × 10 ⁶	0.0489	1.304°	11.86y
Saturn	60270	568.5 × 10 ²⁴	10.66h	26.73°	1.433 × 10 ⁹	0.0565	2.485°	29.46y
Uranus	25560	86.83 × 10 ²⁴	17.24h*	97.77°	2.872 × 10 ⁹	0.0457	0.772°	84.01y
Neptune	24760	102.4 × 10 ²⁴	16.11h	28.32°	4.495 × 10 ⁹	0.0113	1.769°	164.8y
(Pluto)	1195	12.5 × 10 ²¹	6.387d	122.5°	5.870 × 10 ⁸	0.2444	17.16°	247.7y

*Retrograde

Astronomical Data

Regions of the Electromagnetic Spectrum	Wavelength (m)	Frequency (Hz)	Photon Energy (J)
Radio	> 1 × 10 ¹	< 3 × 10 ⁹	> 2 × 10 ²⁴
Microwave	1 × 10 ⁻³ - 1 × 10 ⁻¹	3 × 10 ⁹ - 3 × 10 ¹¹	2 × 10 ⁻²⁴ - 2 × 10 ⁻²²
Infrared	7 × 10 ⁻⁷ - 1 × 10 ⁻³	3 × 10 ¹¹ - 4 × 10 ¹⁴	2 × 10 ⁻²² - 3 × 10 ⁻¹⁹
Optical	4 × 10 ⁻⁷ - 7 × 10 ⁻⁷	4 × 10 ¹⁴ - 7.5 × 10 ¹⁴	3 × 10 ⁻¹⁹ - 5 × 10 ⁻¹⁹
UV	1 × 10 ⁻⁸ - 4 × 10 ⁻⁷	7.5 × 10 ¹⁴ - 3 × 10 ¹⁶	5 × 10 ⁻¹⁹ - 2 × 10 ⁻¹⁷
X-ray	1 × 10 ⁻¹¹ - 1 × 10 ⁻⁸	3 × 10 ¹⁶ - 3 × 10 ¹⁹	2 × 10 ⁻¹⁷ - 2 × 10 ⁻¹⁴
Gamma-ray	< 1 × 10 ⁻¹¹	> 3 × 10 ¹⁹	2 × 10 ⁻¹⁴

geostationary orbit altitude	b_{geo}	35786 km
Low Earth Orbit altitude	b_{LEO}	120-3000 km
speed of light	c	299792.458 km/s
universal gas constant (ideal gas)	R	8.3144598 J mol ⁻¹ K ⁻¹
Moon distance from Earth	d	384402 km
Moon radius	R	1737 km
Moon surface acceleration	g	1.625 m/s ²
Mars radius	R	3396 km
Mars surface acceleration	g	3.72 m/s ²
Newtonian constant of gravitation	G	6.673 848 × 10 ⁻¹¹ m ³ kg ⁻¹ s ⁻²
standard acceleration of gravity	g_0	9.806 65 m/s ²
Earth mass	M	5.974 × 10 ²⁴ kg
Earth radius (volumetric mean)	R	6.378 × 10 ⁶ m
solar luminosity	L	3.846 × 10 ²⁶ W
solar mass	M	1.989 × 10 ³⁰ kg
solar radius (volumetric mean)	R	6.960 × 10 ⁸ m
Jupiter mass	M	1.899 × 10 ²⁷ kg
Jupiter radius	R	6.991 × 10 ⁷ m
astronomical unit	$1 UA$	1.495 979 × 10 ¹¹ m
parsec pc	$1 pc$	3.085 678 × 10 ¹⁶ m
Sidereal year	$1 sidereal year$	365.26 days
Sidereal day	$1 sidereal day$	86164.09 s

Frequent Constants

CONSTANTS

ROCKET DYNAMICS

Drag equation $D = qAC_D$	Delta V neglecting drag and gravity $\Delta v = I_{sp} g_0 \ln \frac{m_0}{m_f}$	
$m_e = m_E + m_p + m_{PL}$	$A_e \equiv$ Nozzle exit area	
$m_p = m_E + m_{PL}$	$T \equiv$ Thrust	
$m_E \equiv$ Empty / structural mass	$I_{sp} \equiv$ Specific impulse	
$m_p \equiv$ Propellant mass	$n \equiv$ Mass ratio	
$m_{PL} \equiv$ Payload mass	$\Delta t \equiv$ Burn time	
$\lambda \equiv$ Payload ratio	$\Delta v_D, \Delta v_G \equiv$ Drag loss, gravity loss	
$\varepsilon \equiv$ Structural ratio	$\gamma \equiv$ Flight path angle	
$D, C_D \equiv$ Drag, Drag coefficient	$v_{bo} \equiv$ Burnout speed	
Specific Impuls $I_{sp} = \frac{T}{\dot{m}_e g_0}$	Effective exhaust velocity $c = c_a + \frac{(p_e - p_a) A_e}{\dot{m}_e}$	Thrust $T = \dot{m}_e c$
Mass Ratio $n = \frac{m_0}{m_f}$	Burn Time $\Delta t = \frac{n-1}{n} \frac{I_{sp}}{T m_e g_0}$	Payload Ratio $\lambda = \frac{m_{PL}}{m_E + m_p} = \frac{m_{PL}}{m_0 - m_{PL}}$
Structural Ratio $\varepsilon = \frac{m_E}{m_E + m_p} = \frac{m_E}{m_0 - m_{PL}}$	Delta V $\Delta v = I_{sp} g_0 \ln \frac{m_0}{m_f} - \Delta v_D - \Delta v_G$	
Burnout Speed $v_{bo} = I_{sp} g_0 \ln n = I_{sp} g_0 \ln \frac{1+\lambda}{\varepsilon + \lambda}$	Drag and gravity loss $\Delta v_D = \int_0^{t_f} \frac{D}{m} dt$ $\Delta v_G = \int_0^{t_f} g \sin \gamma dt$	
Optimization procedure to an N- stage vehicle η - Lagrange multiplier	$h = \sum_{i=1}^N [\ln(1 - \varepsilon_i) + \ln n_i - \ln(1 - \varepsilon_i n_i)] - \eta \left(v_{bo} - \sum_{i=1}^N c_i \ln n_i \right)$	

IDEAL GAS RELATIONS

Ideal gas model	$p v = RT$ $u = u(T)$ $h = h(T) = u(T) + RT$
Change in specific internal energy	$u(T_2) - u(T_1) = \int_{T_1}^{T_2} c_v(T) dT$ For constant C_v $u(T_2) - u(T_1) = c_v(T_2 - T_1)$
Change in specific enthalpy	$h(T_2) - h(T_1) = \int_{T_1}^{T_2} c_p(T) dT$ For constant C_p $h(T_2) - h(T_1) = c_p(T_2 - T_1)$

COMPRESSIBLE FLOW IN NOZZLES AND DIFFUSERS

Momentum equation for steady-state one dimension flow	$F = \dot{m}(V_2 - V_1)$
Ideal gas velocity of sound	$c = \sqrt{kRT}$
Mach number	$M = V/c$
Stagnation enthalpy	$h_o = h + V^2/2$
Isentropic flow function relating temperature and stagnation temperature (constant k)	$\frac{T_o}{T} = 1 + \frac{k-1}{2} M^2$
Isentropic flow function relating pressure and stagnation pressure (constant k)	$\frac{p_o}{p} = \left(\frac{T_o}{T} \right)^{k/(k-1)} = \left(1 + \frac{k-1}{2} M^2 \right)^{k/(k-1)}$

VALISPACE

EMPOWERING ENGINEERS

SPACE
CHEAT SHEET

contact-us@valispace.com
www.valispace.com

$i \equiv$ inertial frame, $s \equiv$ reference frame

$\omega^s \equiv$ angular velocity of the s frame with respect to the i frame

$r \equiv$ position vector in s frame, $v \equiv$ velocity, $a \equiv$ acceleration

$R \equiv$ position vector of the origin of s-frame

$${}^i v = \frac{d}{dt} r = \frac{d}{dt} r + \omega^{st} \times r$$

$$\frac{d^2}{dt^2} (R+r) = \frac{d^2}{dt^2} R + \frac{d^2}{dt^2} r + 2\omega^{st} \times \frac{d}{dt} r + \left(\frac{d}{dt} \omega^{st} \right) \times r + \omega^{st} \times (\omega^{st} \times r)$$

Polar Coordinates

Unit Vectors $e_r = \cos \theta i + \sin \theta j$
 $e_\theta = -\sin \theta i + \cos \theta j$

Kinematic Equations $r = r e_r \quad v = \dot{r} e_r + r \dot{\theta} e_\theta$
 $a = (\ddot{r} - r \dot{\theta}^2) e_r + (r \ddot{\theta} + 2\dot{r} \dot{\theta}) e_\theta$

Motion Equations $F_r = m a_r = m (\ddot{r} - r \dot{\theta}^2)$
 $F_\theta = m a_\theta = m (r \ddot{\theta} + 2\dot{r} \dot{\theta})$

Cylindrical Coordinates

Unit Vectors $e_r = \cos \theta i + \sin \theta j$
 $e_\theta = -\sin \theta i + \cos \theta j$

Kinematic Equations $r = r e_r + z k \quad v = \dot{r} e_r + r \dot{\theta} e_\theta + \dot{z} k$
 $a = (\ddot{r} - r \dot{\theta}^2) e_r + (r \ddot{\theta} + 2\dot{r} \dot{\theta}) e_\theta + \ddot{z} k$

Motion Equations $F_r = m a_r = m (\ddot{r} - r \dot{\theta}^2)$

$F_\theta = m a_\theta = m (r \ddot{\theta} + 2\dot{r} \dot{\theta})$

$F_z = m a_z = m \ddot{z}$

Spherical Coordinates

Unit Vectors $e_r = \cos \theta \cos \phi i + \sin \theta \cos \phi j + \sin \phi k$
 $e_\theta = -\sin \theta i + \cos \theta j$
 $e_\phi = -\cos \theta \sin \phi i - \sin \theta \sin \phi j + \cos \phi k$

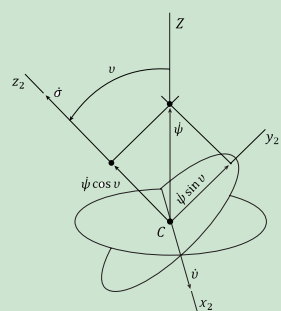
Kinematic Equations $r = r e_r$
 $v = \dot{r} e_r + r \dot{\theta} \cos \phi e_\theta + r \dot{\phi} e_\phi$
 $a = (\ddot{r} - r \dot{\theta}^2 \cos^2 \phi - r \dot{\phi}^2) e_r + (2\dot{r} \dot{\theta} \cos \phi + r \ddot{\theta} \cos \phi - 2r \dot{\theta} \dot{\phi} \sin \phi) e_\theta + (2\dot{r} \dot{\phi} + r \dot{\phi}^2 \sin \phi \cos \phi + r \ddot{\phi}) e_\phi$

Equations of Motion $F_r = m a_r = m (\ddot{r} - r \dot{\theta}^2 \cos^2 \phi - r \dot{\phi}^2)$
 $F_\theta = m a_\theta = m (2\dot{r} \dot{\theta} \cos \phi + r \ddot{\theta} \cos \phi - 2r \dot{\theta} \dot{\phi} \sin \phi)$
 $F_\phi = m a_\phi = m (2\dot{r} \dot{\phi} + r \dot{\phi}^2 \sin \phi \cos \phi + r \ddot{\phi})$

SATELLITE ATTITUDE DYNAMICS

Dual-Spin Spacecraft - Axisymmetric Body

δ - spin, ψ - precession, ν - nutation



Euler's equations for torque-free motion

$$\begin{bmatrix} H_x \\ H_y \\ H_z \end{bmatrix} = \begin{bmatrix} A & 0 & 0 \\ 0 & B & 0 \\ 0 & 0 & C \end{bmatrix} \begin{bmatrix} \omega_x \\ \omega_y \\ \omega_z \end{bmatrix} = \begin{bmatrix} A\omega_x \\ B\omega_y \\ C\omega_z \end{bmatrix}$$

Angular Position of the body in the spinning reference frame $\{\hat{\Omega}\}_{x_2 y_2 z_2 \equiv x y z} = \vec{\psi} + \vec{\nu} = \begin{bmatrix} \dot{\psi} \sin \nu \\ \dot{\psi} \cos \nu \\ \dot{\nu} \end{bmatrix}$

Angular velocity in the spinning reference frame $\{\hat{\omega}\}_{x_2 y_2 z_2 \equiv x y z} = \{\hat{\Omega}\}_{x_2 y_2 z_2 \equiv x y z} + \vec{\sigma} = \begin{bmatrix} \dot{\psi} \sin \nu \\ \dot{\psi} \cos \nu + \dot{\sigma} \\ \dot{\nu} \end{bmatrix}$

Angular moment components $(\vec{H}_c)_i = \begin{bmatrix} A\dot{\nu} \\ A\dot{\psi} \sin \nu \\ C(\dot{\psi} \cos \nu + \dot{\sigma}) \end{bmatrix}$

γ - wobble angle

Relations between precession and spin

$$\dot{\psi} = \frac{C}{(A-C) \cos \nu} \dot{\sigma} \quad \tan \gamma = \frac{\omega_y}{\omega_z} = \frac{\dot{\psi} \sin \nu}{\dot{\psi} \cos \nu + \dot{\sigma}} \quad \tan \gamma = \frac{C}{A} \tan \nu$$

Central Force

$H \equiv$ Angular momentum $v \equiv$ velocity
 $b \equiv$ specific angular momentum $E \equiv$ Total Energy
 $\frac{dA}{dt} \equiv$ areal velocity $U \equiv$ Potential Energy
 $r \equiv$ position
 $\mu \equiv$ gravitational parameter

Formal solution of the central force $t - t_0 = \int_{r_0}^r \frac{dr}{\pm \sqrt{\frac{2}{m} (E - U(r)) - \frac{h^2}{r^2}}}$

Kepler's 2nd law $\frac{dA}{dt} = \frac{H_0}{2m} = \frac{h}{2} = \text{constant}$

Conservation of angular momentum $\vec{H}_0 = \vec{r} \times m\vec{v} = \text{Cte}$

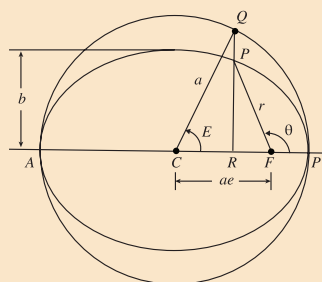
Specific angular momentum $h = r^2 \dot{\theta}$

Central force $\vec{F} = F \frac{\vec{r}}{r}$

Gravitational central force $F = -\frac{GMm}{r^2} e_r; f = -\frac{\mu}{r^2}$

Ellipse Properties

$e \equiv$ eccentricity $\epsilon \equiv$ specific energy
 $a \equiv$ semimajor axis $v_r \equiv$ radial velocity
 $b \equiv$ semiminor axis $\gamma \equiv$ flight path angle
 $p \equiv$ semi latus rectum / parameter $T \equiv$ orbital period
 $r_p \equiv$ periapsis $n \equiv$ mean motion
 $r_a \equiv$ apoapsis $M \equiv$ mean anomaly
 $F \equiv$ Focus, $C \equiv$ Center $E \equiv$ eccentric anomaly
 $c \equiv$ linear eccentricity $\theta \equiv$ true anomaly
 $v_{esc} \equiv$ escape velocity



$$2a = r_p + r_a \quad b = a\sqrt{1 - e^2} \quad \frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

$$r_p = a(1 - e) = \frac{p}{1 + e} \quad \vec{CF} = c = ae \quad \text{Area} = abn$$

$$r_a = a(1 + e) \quad e = \sqrt{1 - \left(\frac{b}{a}\right)^2} \quad p = a(1 - e^2) = \frac{h^2}{\mu}$$

Elliptical Orbits

Gravitational Parameter $\mu = GM$

Specific energy $\epsilon = -\frac{\mu}{2a}$

eccentricity $e = \sqrt{1 + 2\epsilon \left(\frac{h}{\mu}\right)^2}$

Mean motion $n = \frac{2\pi}{T} = \sqrt{\frac{\mu}{a^3}}$

Radial velocity $v_r = \frac{\mu}{h} e \sin \theta$

Escape velocity $v_{esc} = \sqrt{\frac{2\mu}{r}}$

Orbital period $T = \frac{2\pi}{\mu} a^{3/2}$

Kepler's 3rd law $\left(\frac{T_2}{T_1}\right)^2 = \left(\frac{a_2}{a_1}\right)^3$

Eccentricity vector $\vec{e} = \frac{1}{\mu} \left(\vec{v} \times \vec{h} - \frac{\mu \vec{r}}{r} \right)$

Orbit equation Kepler's 1st law $r = \frac{h^2}{\mu} \frac{1}{1 + e \cos \theta} = \frac{p}{1 + e \cos \theta}$

vis - viva equation $\frac{v^2}{2} - \frac{\mu}{r} = \epsilon = cte; v = \sqrt{2\mu \left(\frac{1}{r} - \frac{1}{2a}\right)}$

Kepler's Equation $M = n(t - t_0) = E - e \sin E$

E as function of θ $\tan E = \frac{\sqrt{1 - e^2} \sin \theta}{e + \cos \theta}, \tan \frac{E}{2} = \sqrt{\frac{1 - e}{1 + e}} \tan \frac{\theta}{2}$

Circular Orbit

Orbit equation $r = \frac{h^2}{\mu}$ velocity $v = \sqrt{\frac{\mu}{r}}$

Time versus true anomaly $t = \frac{\theta}{2\pi} T$ specific energy $\epsilon = -\frac{\mu}{2r}$

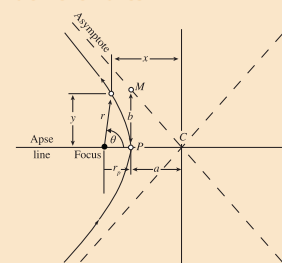
Orbital period $T = \frac{2\pi}{\mu} r^{3/2}$

Parabolic orbits

Parabolic mean anomaly $M_p = \frac{\mu^2}{h^3} t$

Barker's equation $2 \sqrt{\frac{\mu}{p^3}} (t - t_0) = \tan \frac{\theta}{2} + \frac{1}{3} \tan^3 \frac{\theta}{2}$

Hyperbolic orbits



$$\sinh F = \frac{y}{b}$$

Hyperbolic mean anomaly $M_h = \frac{\mu^2}{h^3} (e^2 - 1)^{3/2} t$

Kepler's equation for the hyperbola $\sqrt{\frac{\mu}{-a^3}} (t - t_0) = e \sinh F - F$

True anomaly in terms of F $\cos \theta = \frac{e - \cosh F}{e \cosh F - 1}$

Orbital Elements

$b \equiv$ specific angular momentum $\omega \equiv$ argument of perigee
 $i \equiv$ inclination $\theta \equiv$ true anomaly
 $\Omega \equiv$ right ascension (RA) of the ascending node $\{i, \Omega, \varpi, a, e, T_0\}$
 $e \equiv$ eccentricity $\vec{n} - \text{nodes line}$

$$a = -\frac{\mu}{v_0^2 - \frac{\mu}{r_0}} \quad \vec{e} = \frac{1}{\mu} \left[\vec{v}_0 \times (\vec{r}_0 \times \vec{v}_0) - \mu \frac{\vec{r}_0}{r_0} \right] \quad \cos \theta_0 = \frac{\vec{e} \cdot \vec{r}_0}{e r_0}$$

$$\vec{n} = \frac{\vec{e}_z \times \vec{h}}{|\vec{e}_z \times \vec{h}|} \quad \vec{n} = \cos \Omega \vec{e}_x + \sin \Omega \vec{e}_y \quad \cos i = \frac{\vec{e}_z \cdot \vec{h}}{|\vec{h}|}$$

$$\cos \varpi = \frac{\vec{n} \cdot \vec{e}}{|\vec{e}|} \quad \tan \theta_0 = \frac{\left(\frac{r_0 v_0}{\mu}\right) \sin \gamma_0 \cos \gamma_0}{\left(\frac{r_0 v_0}{\mu}\right) \cos^2 \gamma_0 - 1} \quad e^2 = \left(\frac{r_0 v_0}{\mu} - 1\right)^2 \cos^2 \gamma_0 + \sin^2 \gamma_0$$

Effects of the Earth's oblateness

$$\dot{\Omega} = -\frac{3nJ_2 R_\oplus^2}{2a^2(1 - e^2)^2} \cos i \quad \dot{\varpi} = \frac{3nJ_2 R_\oplus^2}{2a^2(1 - e^2)^2} \left(2 - \frac{5}{2} \sin^2 i\right)$$

Oblateness and second zonal harmonics

Planet	Oblateness	J_2
Mercury	0.000	60×10^{-6}
Venus	0.000	4.458×10^{-6}
Earth	0.003353	1.08263×10^{-3}
Mars	0.00648	1.96045×10^{-3}
Jupiter	0.06487	14.736×10^{-3}
Saturn	0.09796	16.298×10^{-3}
Uranus	0.02293	3.34343×10^{-3}
Neptune	0.01708	3.411×10^{-3}
(Moon)	0.0012	202.7×10^{-6}

3 bodies

Jacobi integral

$$\frac{1}{2} v^2 - \frac{1}{2} (x^2 + y^2) - \frac{1 - \mu}{r_1} - \frac{\mu}{r_2} = C$$

Orbital Maneuvers

Hohmann Transfer

ϕ - Phase angle $2a_H = r_1 + r_2$

$$\Delta v_1 = \sqrt{2\mu \left(\frac{1}{r_1} - \frac{1}{2a_H}\right)} - \sqrt{\frac{\mu}{r_1}} \quad \Delta v_2 = \sqrt{\frac{\mu}{r_2}} - \sqrt{2\mu \left(\frac{1}{r_2} - \frac{1}{2a_H}\right)}$$

time during transfer $-t_t = \frac{T_H}{2} = \pi \sqrt{\frac{a_H^3}{\mu}}$

$$T_{syn} = \frac{T_1 T_2}{|T_1 - T_2|} \quad \phi = \phi_f + (n_2 - n_1)t \quad t_{wait} = \frac{-2\phi_f \pm 2\pi N}{n_2 - n_1}$$